

Overview of the Milky Way Galaxy

Luminous (stellar) Mass:

$$3 \times 10^{10} M_{\text{sun}}$$

Dynamical Mass (total)

$$2 \times 10^{12} M_{\text{sun}}$$

Interstellar Mass:

$$\text{Neutral (H I)} \quad 3 \times 10^9 M_{\text{sun}}$$

$$\text{Ionized (H II)} \quad 1 \times 10^9 M_{\text{sun}}$$

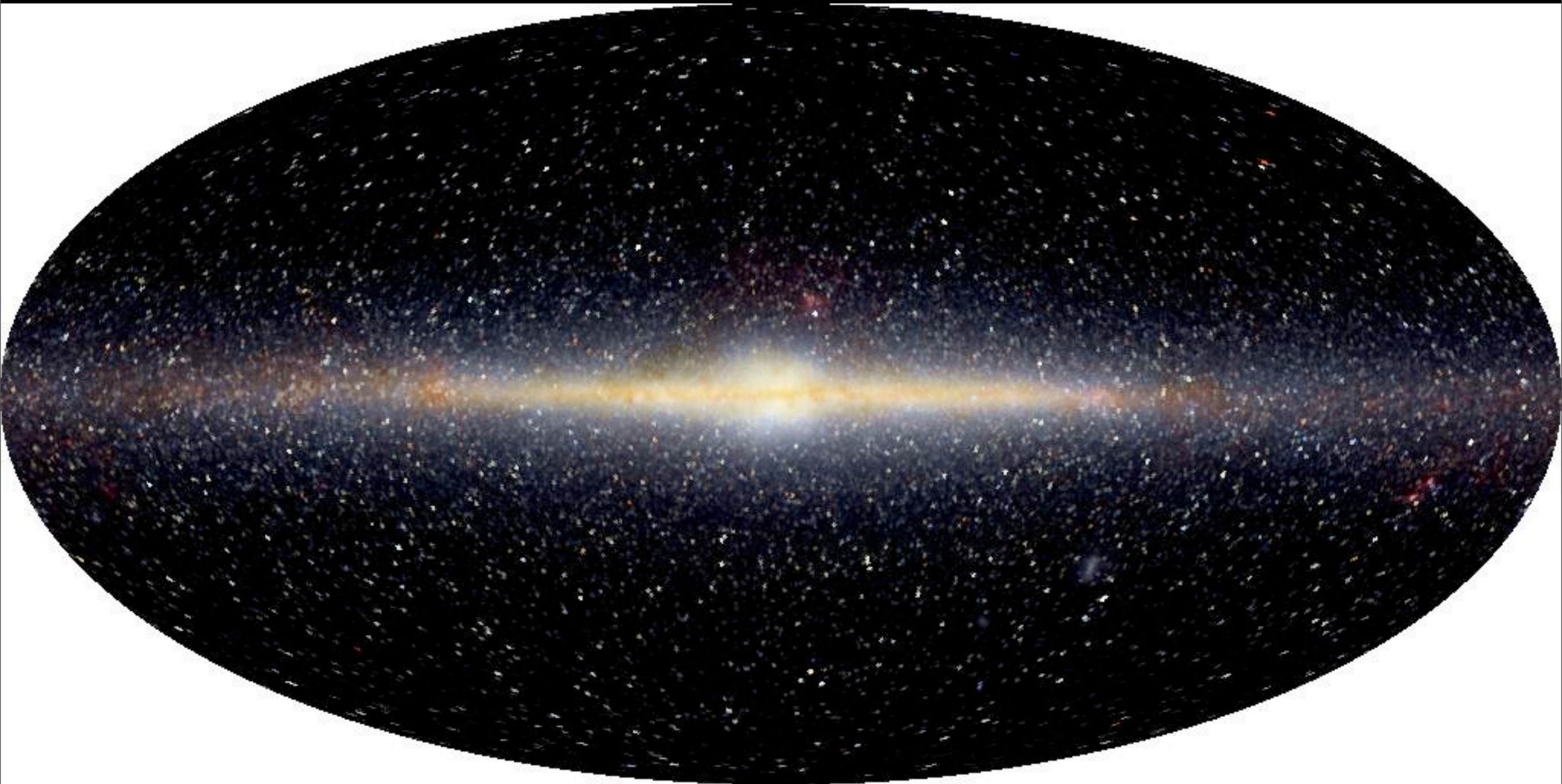
$$\text{Molecular (H}_2\text{)} \quad \underline{1 \times 10^9 M_{\text{sun}}}$$

$$7 \times 10^9 M_{\text{sun}}$$

→ Total incl Helium (36%)



Milky Way (infrared image)



Milky Way Galactic Center (Sagittarius)

(Spitzer Space Telescope)



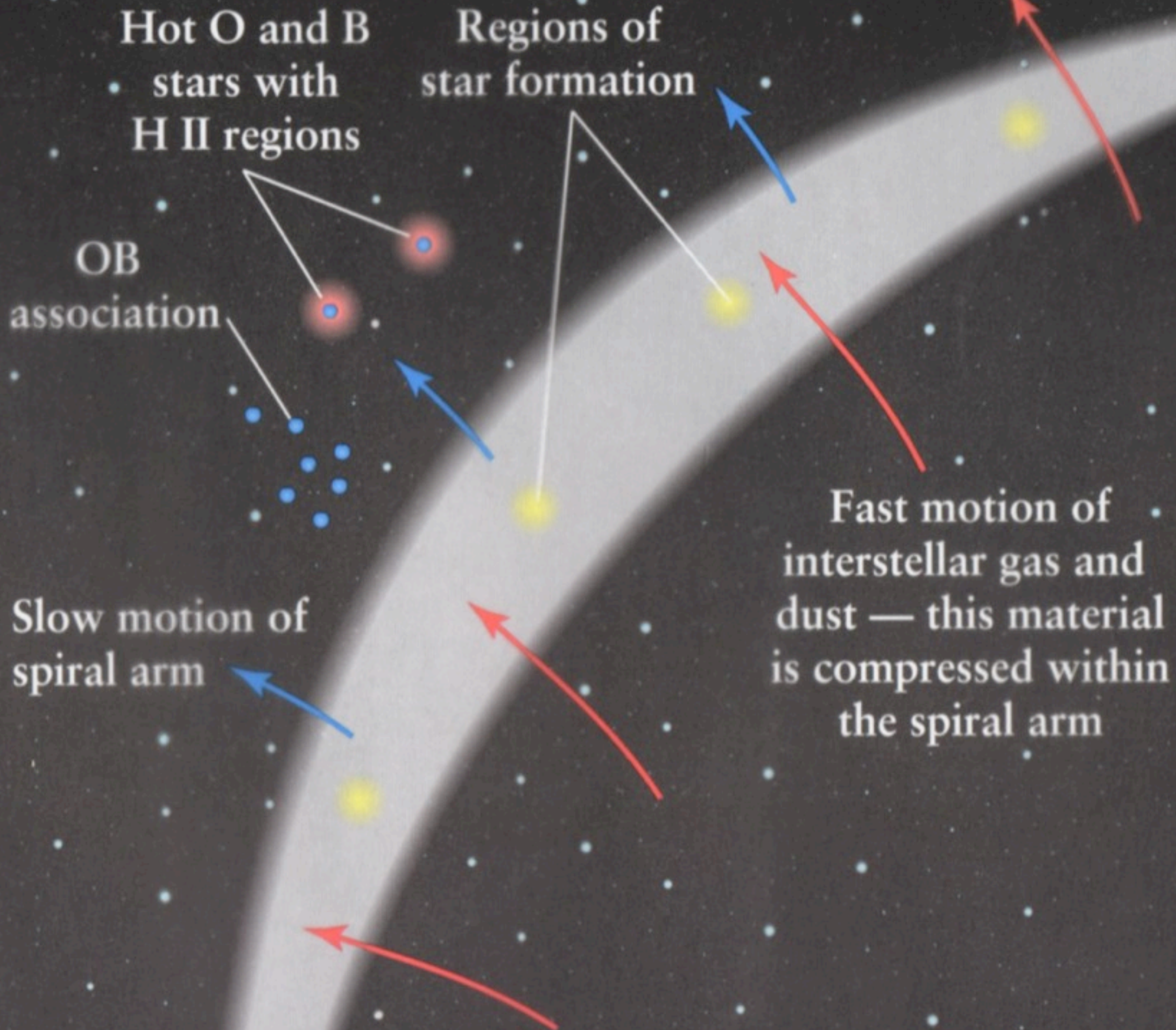
A Spitzer Space Telescope image of the Milky Way Galactic Center, showing a dense field of stars and interstellar dust. The dust is primarily reddish-brown, while the stars appear as white and yellow points of light. A white arrow points to a bright, compact source at the center. A red double-headed arrow at the bottom indicates a scale.

900 ltyrs at $R_0 = 8.4$ kpc

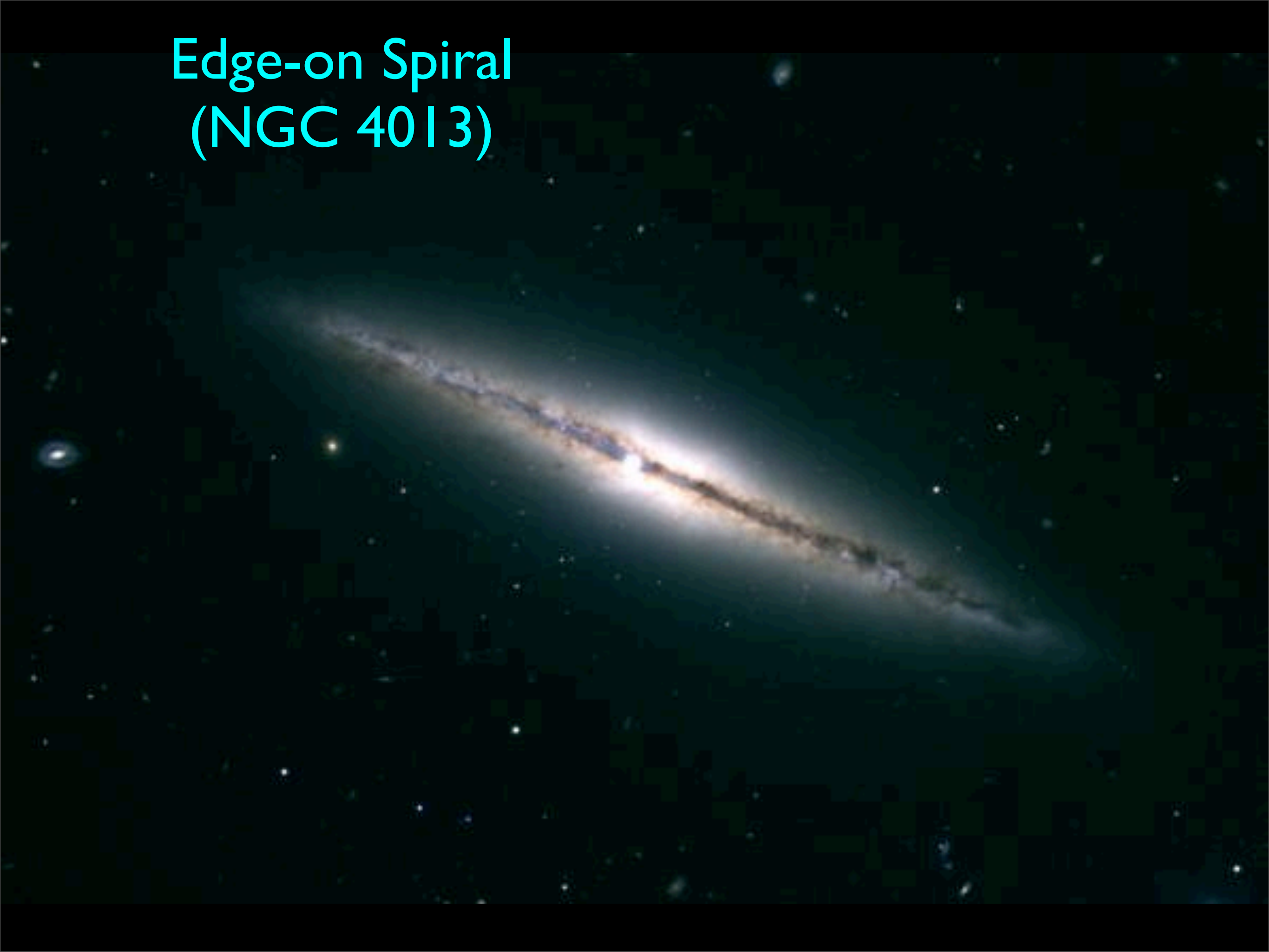
Face-on View

Rotation





Edge-on Spiral (NGC 4013)



Another Spiral Galaxy



Barred Spiral (Hubble image)



NGC 1365

(largest barred spiral)



200,000 light yrs across
(taken by ESA/La Silla
1.5m telescope)



Young stars form in
spiral arms of galaxies

Interstellar gas clouds
are organized and
compacted by “spiral
density waves”

Then, gravity
takes over



LMC

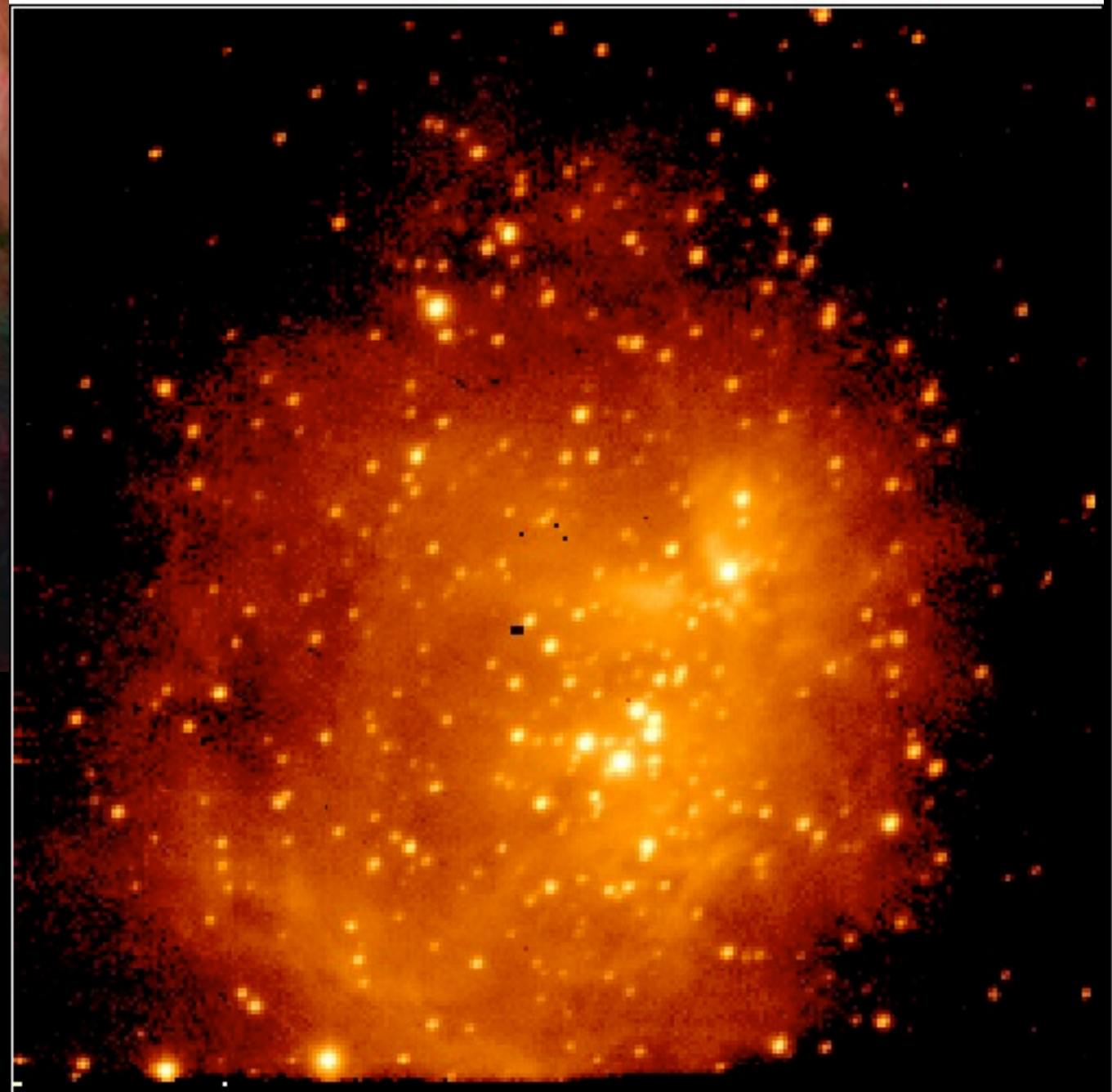
Highlighting
Supernova
Remnants

$d \approx 50$ kpc

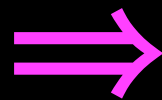
Orion Nebula and Trapezium Stars



Orion Trapezium, IR K-Band, 5.5' x 5.5' FOV



Orion Star Cluster
(infrared image)



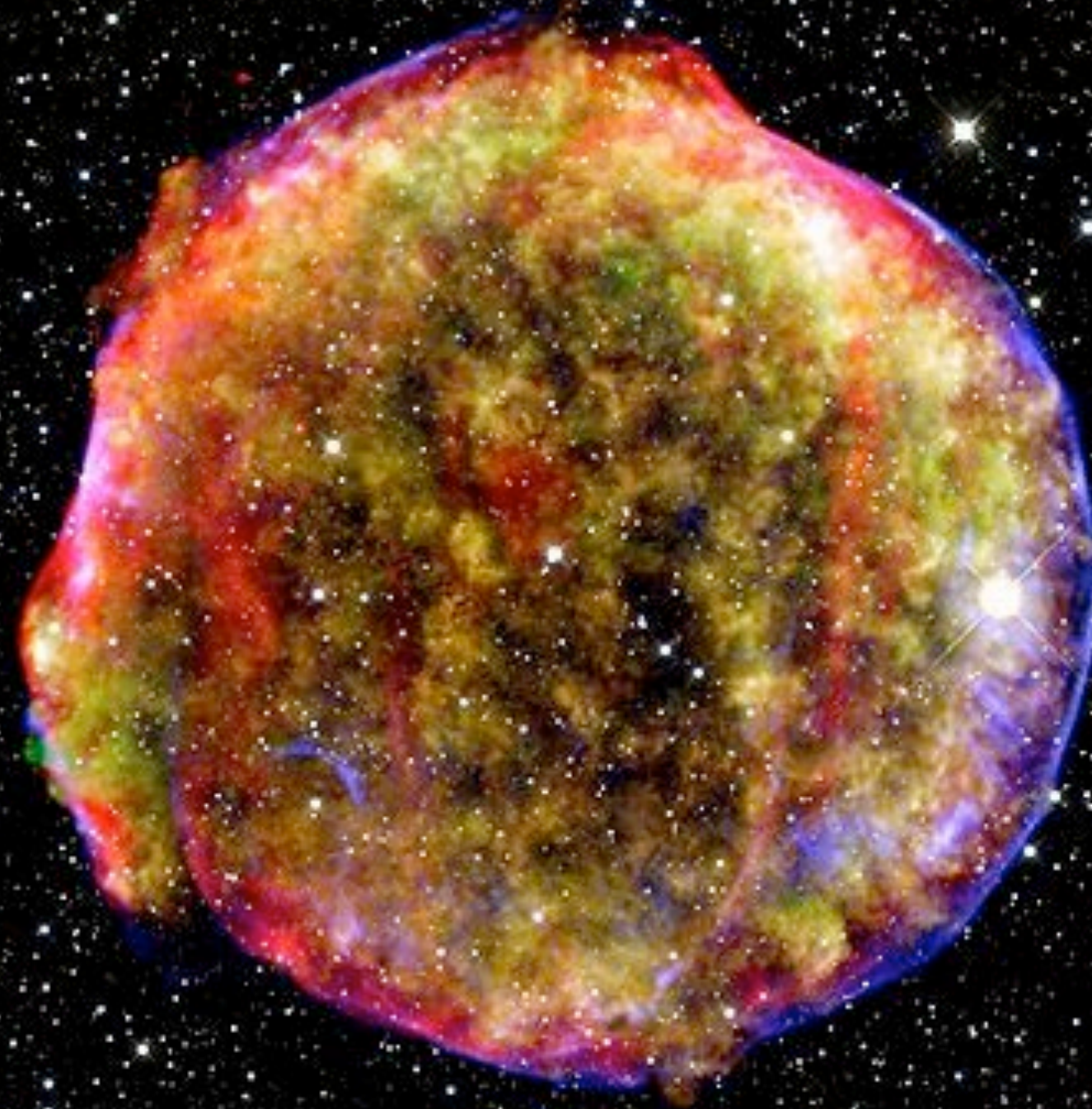
Dying Stars (Planetary Nebulae with White Dwarf)



NGC 2392 • “Eskimo” Nebula **HST • WFPC2**
NASA, A. Fruchter and the ERO Team (STScI) • STScI-PRC00-07



Tycho's SNR (1572)



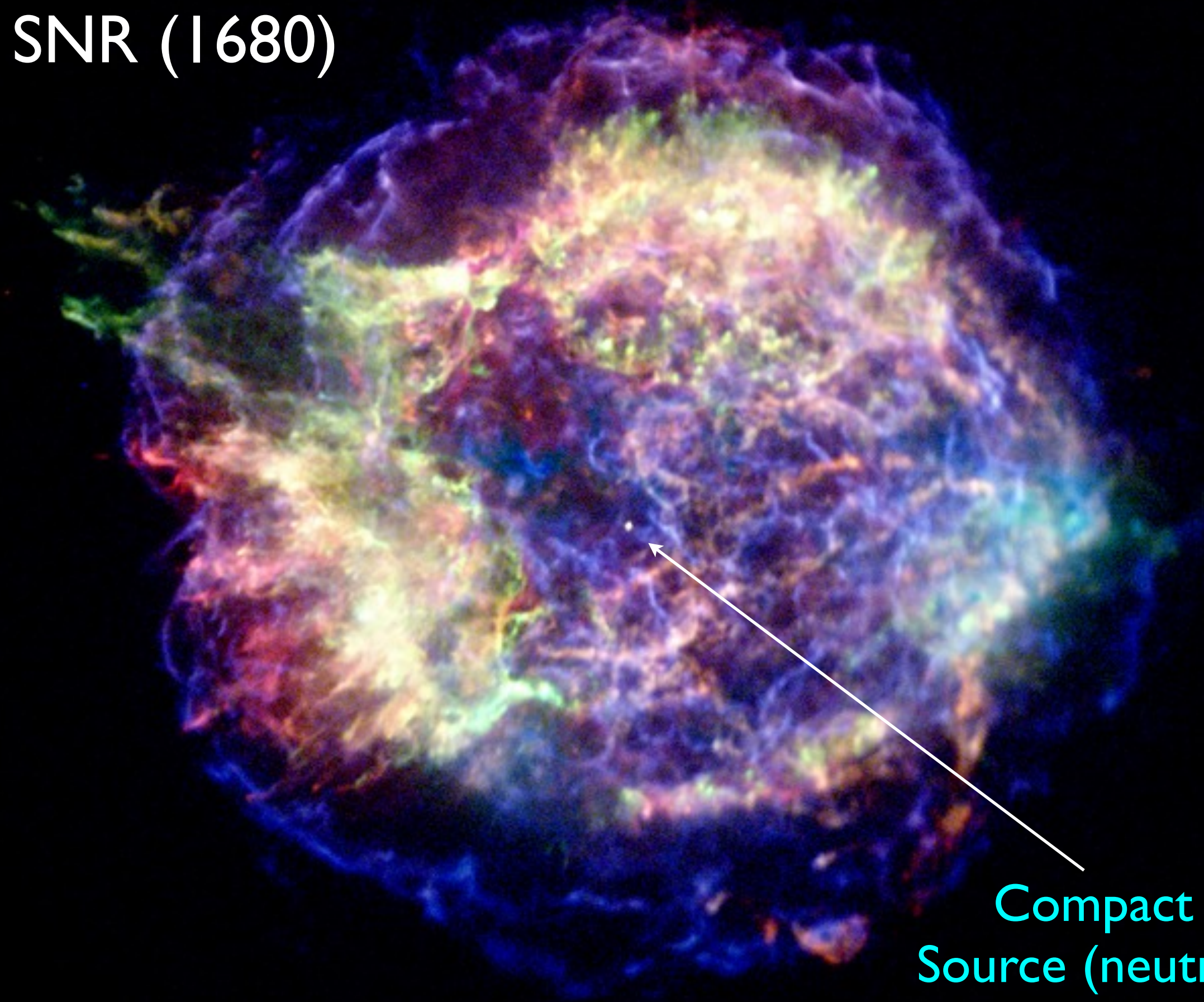
7500 light
years away



Tycho Brahe
1546 - 1601

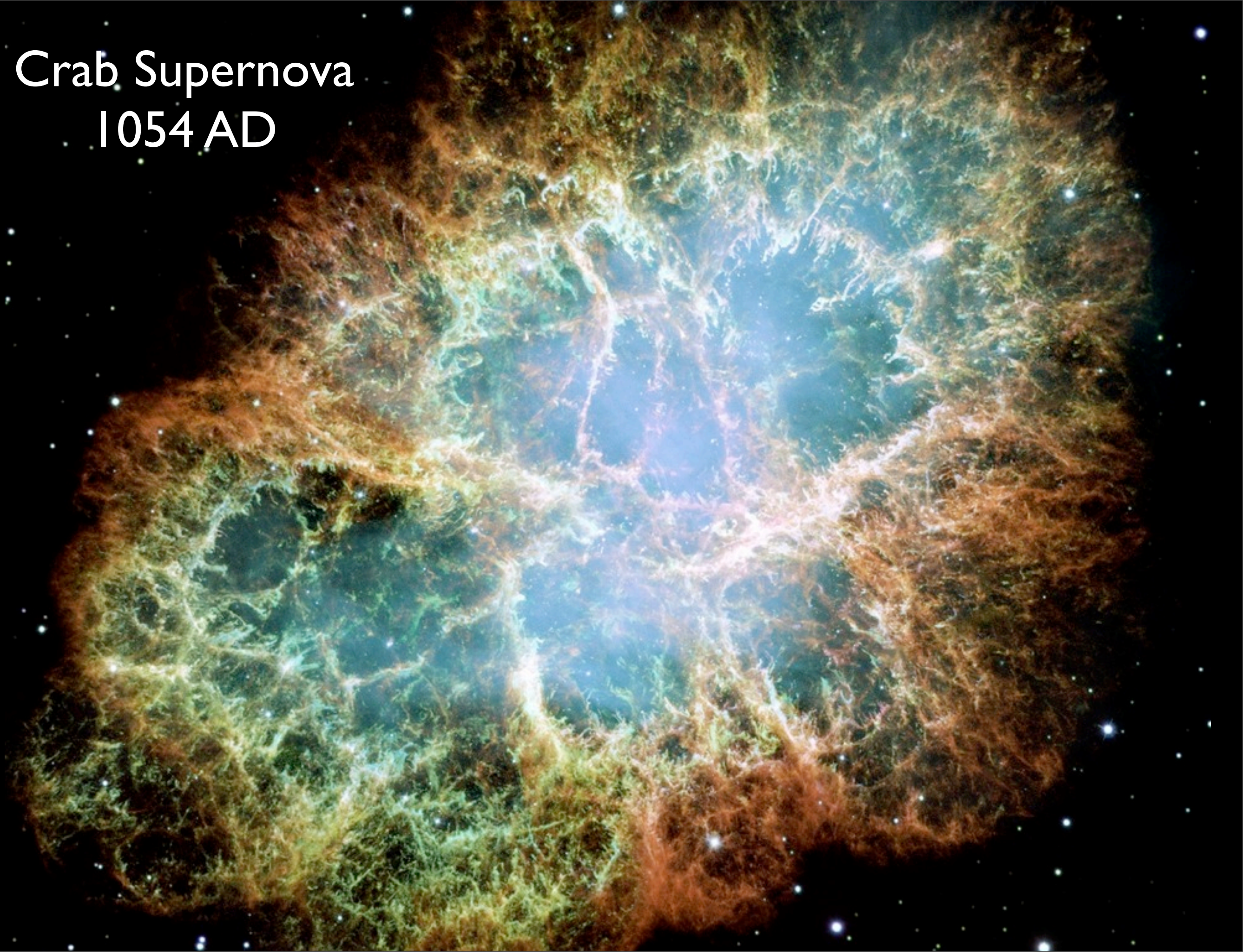
Cassiopeia A

SNR (I 680)



Compact X-ray
Source (neutron star?)

Crab Supernova 1054 AD



S-147 in
Taurus

Older SNR
40,000 yrs ?

